

Reflection and Refraction

Maxwell's equations lead to definition for the **velocity of electromagnetic radiation in a vacuum**:

$$c = \frac{1}{\sqrt{\epsilon_0 \mu_0}}$$

where

ϵ_0 is the **permittivity** of free-space ($8.854 \times 10^{-12} \text{ C}^2 \cdot \text{N}^{-1} \cdot \text{m}^{-2}$)

μ_0 is the **permeability** of free-space ($4 \times 10^{-7} \text{ kg} \cdot \text{m} \cdot \text{C}^{-2}$)

In a medium, velocity is reduced

$$v = \frac{1}{\sqrt{\mu}}$$

The ratio of the velocity in a medium to free-space is **refractive index**

$$= \frac{c}{v} = \sqrt{\frac{\mu}{\epsilon_0 \mu_0}} > 1.00 \text{ in a medium}$$

- **varies with wavelength**
 - usually increases with frequency (called **normal dispersion**)
 - decreases with frequency in region of absorption (called **anomalous dispersion**)

(nm)	
351	1.539
458	1.525
486	1.522
532	1.519
644	1.515
830	1.510

Important:

Frequency of radiation is fixed by source. Hence, wavelength of radiation in a medium must increase

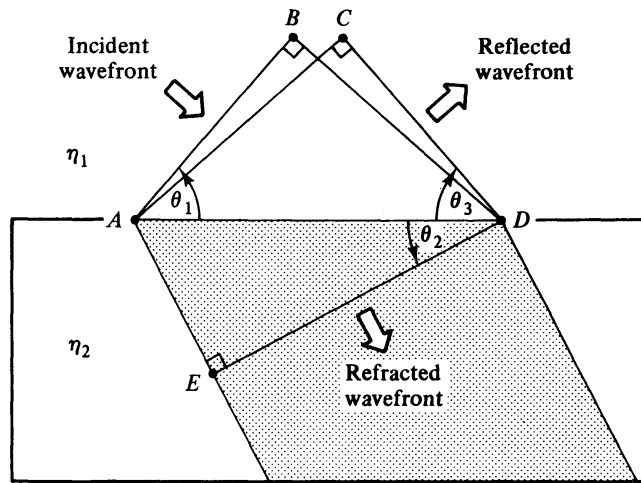
$$\lambda_{\text{medium}} = \frac{c}{\nu} \quad \text{since } \nu = \text{constant}$$

medium > vacuum

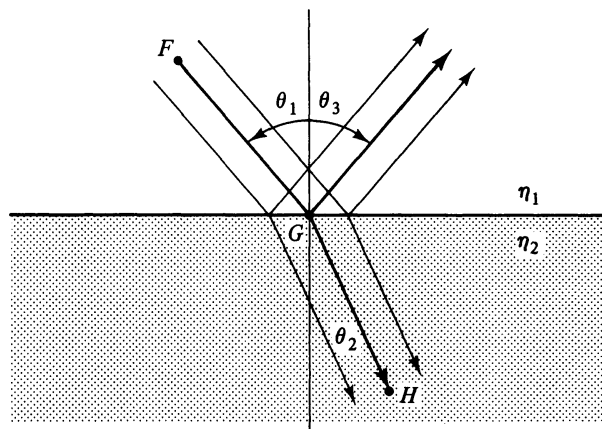
When a wave passes from medium with refractive index n_1 , to medium of refractive index n_2 , we can write

$$\frac{\lambda_2}{\lambda_1} = \frac{c}{\nu} \times \frac{\nu}{c} = \frac{1}{n_2/n_1}$$

Based on wave representation of electromagnetic radiation and geometry, we can quickly deduce the **angle of reflection**:



(a)



(b)

$i = 3$ Law of specular reflectance

The refracted beam does not travel at same velocity as the incident beam ($v_2 = v_1 \cdot n_1 / n_2$):

- first part of the wavefront to strike the interface is retarded preferentially
- light beam **bends towards the interface normal when $n_2 > n_1$**

$$n_1 \sin \theta_1 = n_2 \sin \theta_2 \quad \text{Snell's law of refraction}$$

- **no refraction** when $\theta_1 = 0^\circ$
- **no transmittance** when $\theta_1 > \theta_c$ (critical angle)

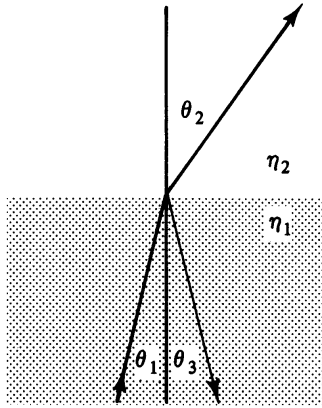
total internal reflection

$$\sin \theta_1 = \frac{n_2}{n_1} \sin \theta_2 \quad \text{Snell's law}$$

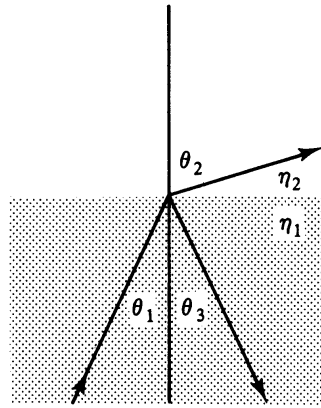
when $\sin \theta_2 = 90^\circ$

$$\theta_i = \theta_c = \sin^{-1} \frac{n_2}{n_1}$$

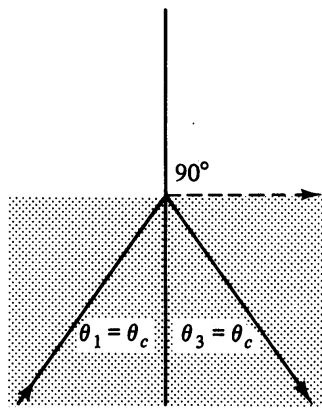
For air/glass $\theta_c = 42^\circ$



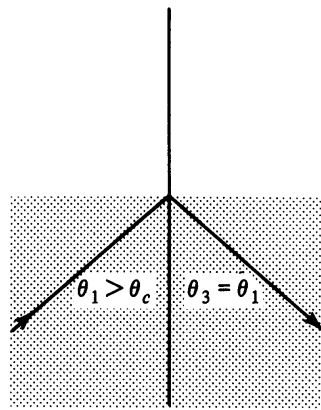
(a)



(b)



(c)



(d)

Fresnel Equations

Reflectance losses occur at all at interfaces

$$R(\theta) + T(\theta) + A(\theta) = 1 \quad \text{Conservation Law}$$

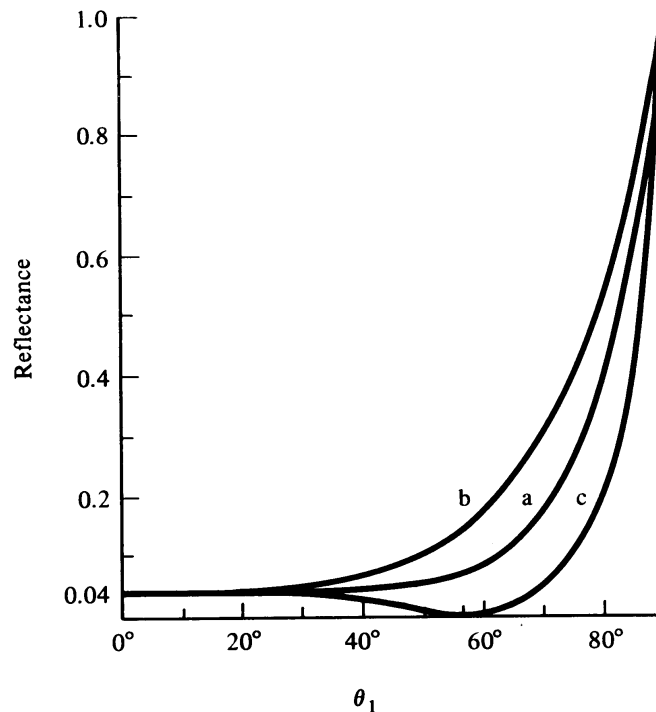
- magnitude increases as the *difference* in the refractive indices increases
- dependent on *incidence angle*

Equation describing the reflectance $R(\theta)$ is the Fresnel equation

$$R(\theta) = \frac{1}{2} \frac{\sin^2(\theta_i - \theta_r)}{\sin^2(\theta_i + \theta_r)} + \frac{\tan^2(\theta_i - \theta_r)}{\tan^2(\theta_i + \theta_r)}$$

- where θ_i is incidence angle and θ_r is refraction angle

For the air/glass at 589 nm, reflectance is about 0.04 or 4 % per interface



- () **constant** for small angles
- () **increases rapidly** at large angles (grazing incidence)

Prisms

Serves several purposes in a spectrometer

- change the **direction** of a beam
- change the **polarization** of a beam
- **split** a beam into two
- **disperse** the beam

A variety of shapes and materials are available to perform these functions.

Dispersing prism

According to Snell's Law,

$$\sin \theta_1 = \frac{n_2}{n_1} \sin \theta_2 \quad \text{Snell's law}$$

there will be **no dispersion** if () is constant

- dispersion in prism occurs because of the **change in refractive index of the prism material as a function of wavelength**
- if prism material exhibits normal dispersion, **higher frequency** (shorter wavelength) light **experiences a higher refractive index** than lower frequency (longer wavelength) light

Light of different wavelengths become divergent and become separated in space

angle between incident and refracted beam is called the **deviation**

The variation in **deviation** with **wavelength** is called the **angular dispersion**

$$D_A = \frac{d}{d} = \frac{d}{d} \underbrace{\frac{d}{d}}_{\text{prism dispersion}}$$

- first term depends on size and shape of the prism and the incidence angle
- second term (prism dispersion) depends on the material of the prism and the wavelength

$$\frac{d}{d} (\text{glass @ 357 nm}) = 1.94 \times 10^{-4} \text{ nm}^{-1}$$

$$\frac{d}{d} (\text{glass @ 825 nm}) = 1.78 \times 10^{-5} \text{ nm}^{-1}$$

Prisms not often used as dispersion elements because of non-constant D_A with wavelength

- produces *non-constant bandwidth*
- means range of 's projected onto exit slit varies with

Electromagnetic radiation

An electromagnetic wave is a transverse wave: electric and magnetic fields perpendicular to the propagation direction

Plane (linearly) polarized beam has constant plane containing the electric and magnetic vectors (often called unpolarized)

The **time-dependent electric field** is

$$E = E_0 \sin(\omega t - kx)$$

where

E_0 is the maximum electric field strength

ω is the angular frequency ($2\pi f$)
 t is time

kx is the (angular) phase

The **angular phase** is ($\phi_0 + 2\pi x/\lambda$) where x is distance and ϕ_0 is the phase at $x=0$

$2/\lambda$ is number of waves per unit length

If two waves maintain the same relative phase difference over

(i) extended period of time

(ii) length

they are said to be coherent

Superposition:

The superposition of two waves states two plane polarized waves can be **algebraically summed** to produce a resultant wave

If waves have same frequency

$$\begin{aligned} E &= E_1 + E_2 \\ &= E_{0,1} \sin(\omega t + \phi_1) + E_{0,2} \sin(\omega t + \phi_2) \end{aligned}$$

Amplitude (intensity) of wave is E^2

$$\begin{aligned} E^2 &= (E_1 + E_2)^2 \\ &= E_1^2 + E_2^2 + 2E_1 E_2 \\ &= E_{0,1}^2 + E_{0,2}^2 + \underbrace{2E_{0,1} E_{0,2} \cos(\phi_2 - \phi_1)}_{\text{interference term}} \end{aligned}$$

If $(\phi_1 - \phi_2) = 0, 2\pi, 4\pi, \dots$

- $\cos(0, 2\pi, 4\pi, \dots) = 1$
- wave amplitude will be reinforced (**constructive interference**)

If $(\phi_1 - \phi_2) = \pi, 3\pi, 5\pi, \dots$

- $\cos(\pi, 3\pi, 5\pi, \dots) = -1$
- wave amplitude will be reduced to zero (**destructive interference**)

Interference can result from difference in pathlength

If the waves initially start out with same phase, the difference in phase, ϕ , due to different paths is

$$\begin{aligned} \phi &= (k_1 - k_2) \\ &= \frac{2\pi}{\lambda} x_1 - \frac{2\pi}{\lambda} x_2 \\ &= \frac{2\pi}{\lambda} (x_1 - x_2) \end{aligned}$$

where

x_1 and x_2 are the lengths to the measurement point from source

$2\pi/\lambda$ is the number of a complete waves per unit length

Thus, when $\phi = 0, 2\pi, \dots$ (an integral number of wavelengths)

$$m \frac{2\pi}{\lambda} = \frac{2\pi}{\lambda} (x_1 - x_2)$$

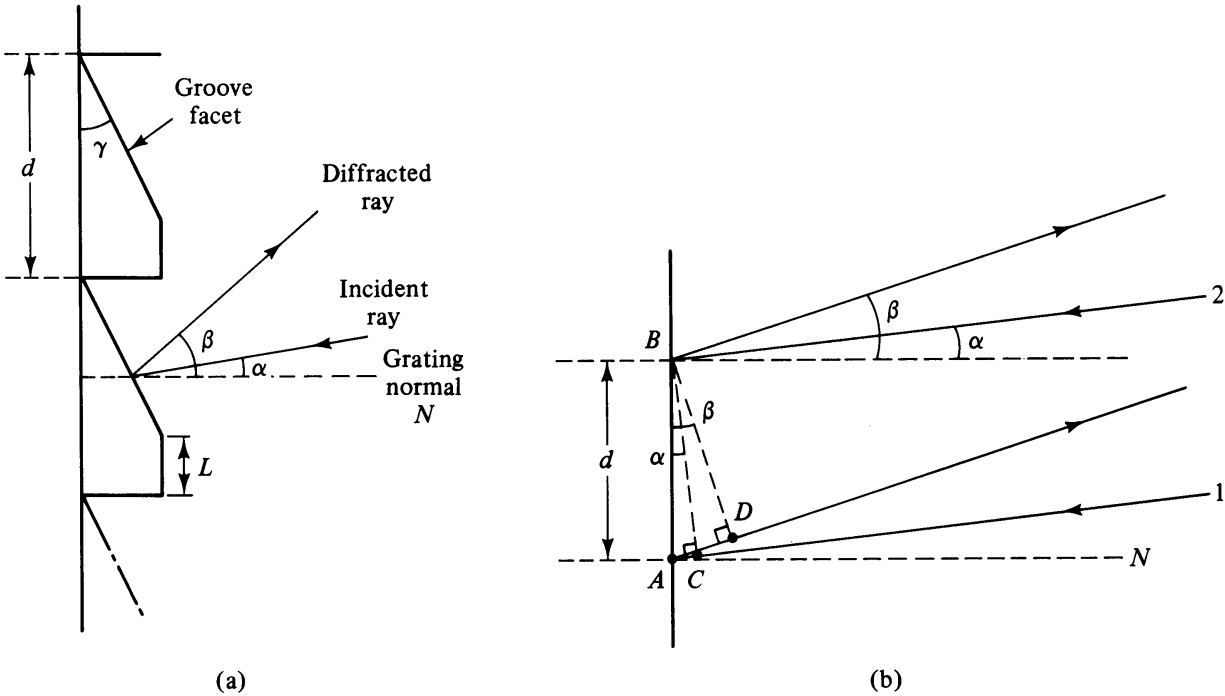
$$m = \frac{x_1 - x_2}{\lambda} \quad \text{constructive interference}$$

when $\phi = \pi, 3\pi, \dots$ (an integral number of wavelengths+1/2)

$$\frac{2m + 1}{2} \frac{2\pi}{\lambda} = \frac{2\pi}{\lambda} (x_1 - x_2) \quad \text{destructive interference}$$

Diffraction (Eschelle) gratings

Parallel grooves etched (blazed) onto reflective surface - **asymmetric in profile**



Incident light striking long facet is reflected in specular direction with respect to the *groove* normal

- light from neighboring grooves travels different distances and so interference occurs in outgoing beam

Note: angles α and β are defined with respect to the *grating* normal, not the *groove* normal

Constructive interference occurs when the pathlength difference is an integral number of wavelengths

- extra pathlength associated with the incident beam is AC

$$AC = d \sin \alpha$$

- extra pathlength associated with the outgoing beam is AD

$$AD = d \sin \beta$$

The total pathlength difference is AC + AD:

$$AC + AD = d(\sin \theta_i + \sin \theta_m)$$

$$m\lambda = d(\sin \theta_i + \sin \theta_m) \quad \text{Grating Formula}$$

(minimum value of d as $\lambda/2$, because the maximum value of $(\sin \theta_i + \sin \theta_m)$ is 2)

The first order ($m = 1$) **diffraction angle** can be calculated for any incidence angle by rearranging the grating formula

$$\frac{m\lambda}{d} = \sin \theta_i + \sin \theta_m$$

$$\sin \theta_m = \frac{m\lambda}{d} - \sin \theta_i$$

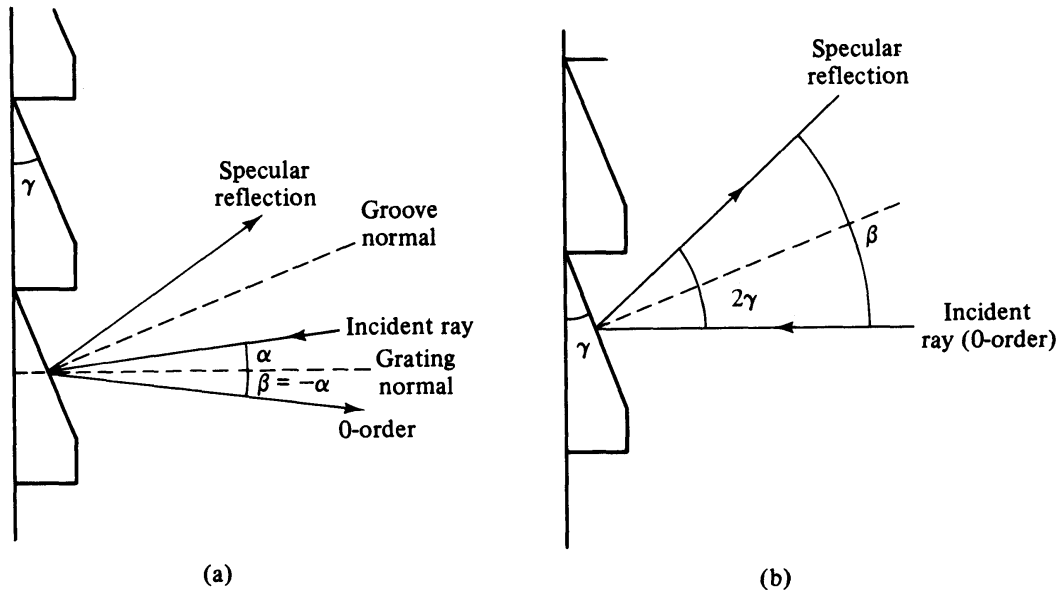
where d is found from the groove spacing

Important

- **diffraction angle** depends on d
- **longer λ 's diffracted more** than shorter ones ($600 \text{ nm} > 500 \text{ nm}$)
- When $m=0$ (zero order), $\sin \theta_m = -\sin \theta_i$ or $\theta_m = -\theta_i$. In this case, **all λ 's are diffracted at the same angle**

If blaze was parallel to the grating plane ($\theta_b = 0^\circ$), the zero order beam would also appear in the specular direction (most of the reflected light not dispersed)

If blaze angle $\neq 0^\circ$, specular and zero-order angles do not correspond and majority of the light is dispersed



In the special case when **incident beam is along the surface normal**, $\alpha = 0$ and **first-order beam** is in **specular** direction

- in this case, β is twice the blaze angle, 2γ . The wavelength at this angle is called the **blaze wavelength**

$$m \lambda_{\text{blaze}} = d(\sin \alpha + \sin \beta)$$

$$\lambda_{\text{blaze}} = d \sin 2\gamma$$

Dispersion

The angular dispersion D_A of the grating can be obtained by differentiating the grating formula with respect to wavelength

For constant incidence angle

$$m = d(\sin \theta_i + \sin \theta_m) \quad \text{Grating Formula}$$

$$\begin{aligned} D_A &= \frac{d}{d\lambda} = \frac{m}{d \cos \theta_m} \quad \text{sin } \theta_i \text{ fixed} \\ &= \frac{d(\sin \theta_i + \sin \theta_m)}{d \cos \theta_m} \\ &= \frac{\sin \theta_i + \sin \theta_m}{\cos \theta_m} \end{aligned}$$

For nearly normal incidence, θ_i is small so θ_m is small, and so $\cos \theta_m$ does not change much with

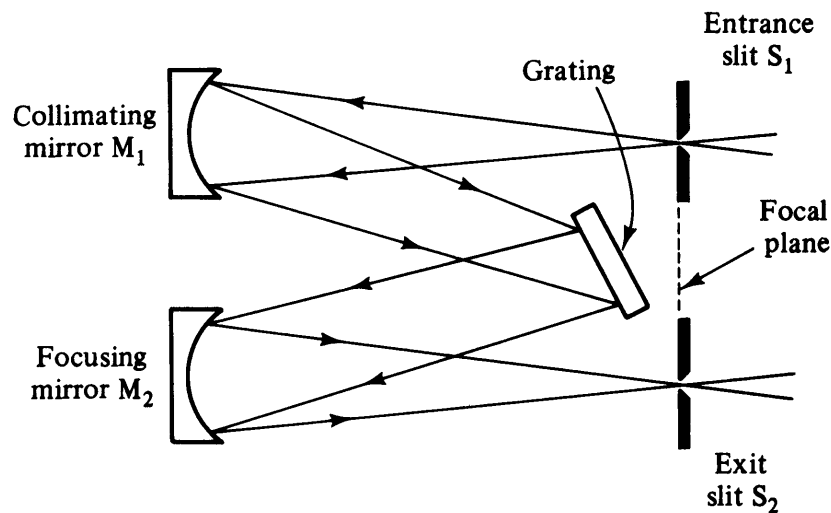
- D_A does not change much with wavelength
- much **better dispersion element** than prism

Monochromators

Comprised of

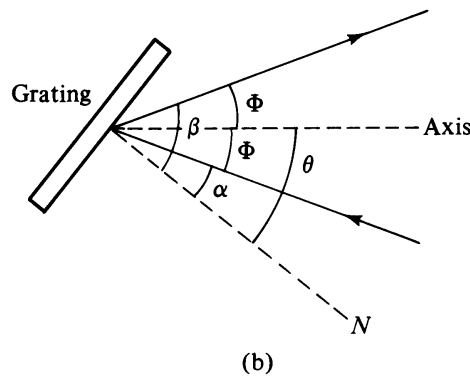
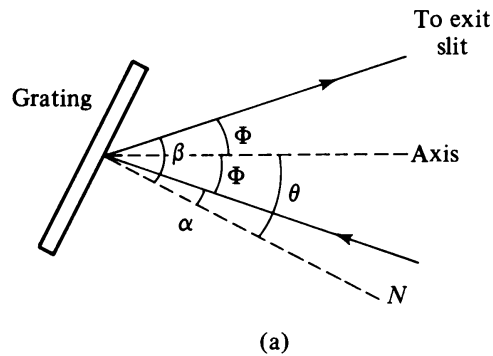
- **dispersive element**
- **image transfer system** (mirrors, lenses and adjustable slits)
 - an image of the entrance slit is transferred to the exit slit after dispersion

One of the most common arrangements is the **Czerny-Turner** monochromator:



Wavelength selection

Wavelength selection is accomplished by **rotating** the grating



Since **angle** between the **entrance slit**, **grating** and **exit slit** is fixed (2 Φ), grating formula can be expressed in terms of the grating rotation angle (between grating normal and optical axis)

Since $\theta = \alpha - \beta$ and $\theta = \alpha + \beta$,

$$m\lambda = d[\sin(\alpha - \beta) + \sin(\alpha + \beta)]$$

$$= 2d \sin \alpha \cos \beta$$

(the trigonometric identity $1/2(\sin(A+B) + \sin(A-B))$ is $\sin A \cdot \cos B$)

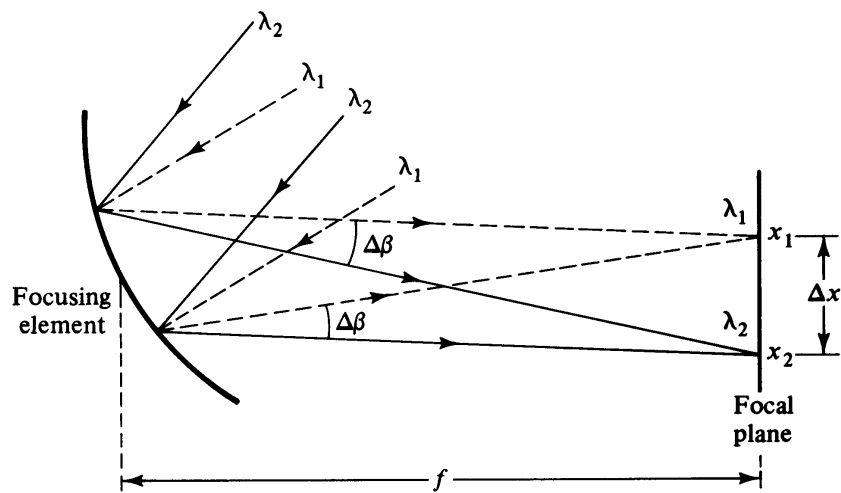
Grating formula now in experimental variables: (the grating rotation angle) and (half-angle between the entrance, grating and exit and slit)

Dispersive characteristics

Already mentioned the angular dispersion (rate of change of diffraction angle with wavelength) for a grating

$$D_A = \frac{d}{d} \quad \text{angular dispersion}$$

However, in monochromator much more interested in dispersion at focal plane (exit slit), defined by the **linear dispersion, D_1** ,



$$D_1 = \frac{dx}{d} \quad \text{linear dispersion}$$

- units of D_1 are $\text{mm}\cdot\text{nm}^{-1}$ or similar

For a Czerny-Turner arrangement, the linear dispersion is:

$$D_1 = f D_A$$

where

f is the **focal length** of the focusing (exit) optic

Sometimes the **inverse linear dispersion**, R_d , is used (units of $\text{nm}\cdot\text{mm}^{-1}$ or similar)

$$R_d = D_1^{-1} = \frac{d}{dx} \quad \text{inverse linear dispersion}$$

$$D_A = \frac{\sin \theta + \sin \theta'}{\cos \theta}$$

$$R_d = (f D_A)^{-1} \\ = \frac{\cos \theta}{f(\sin \theta + \sin \theta')}$$

Spectral bandpass and the slit function

The **spectral bandpass** (nm) is the **half-width** of the range of wavelengths passing through the exit slit

The geometric spectral bandpass

$$s_g = R_d W \quad \text{geometric spectral bandpass}$$

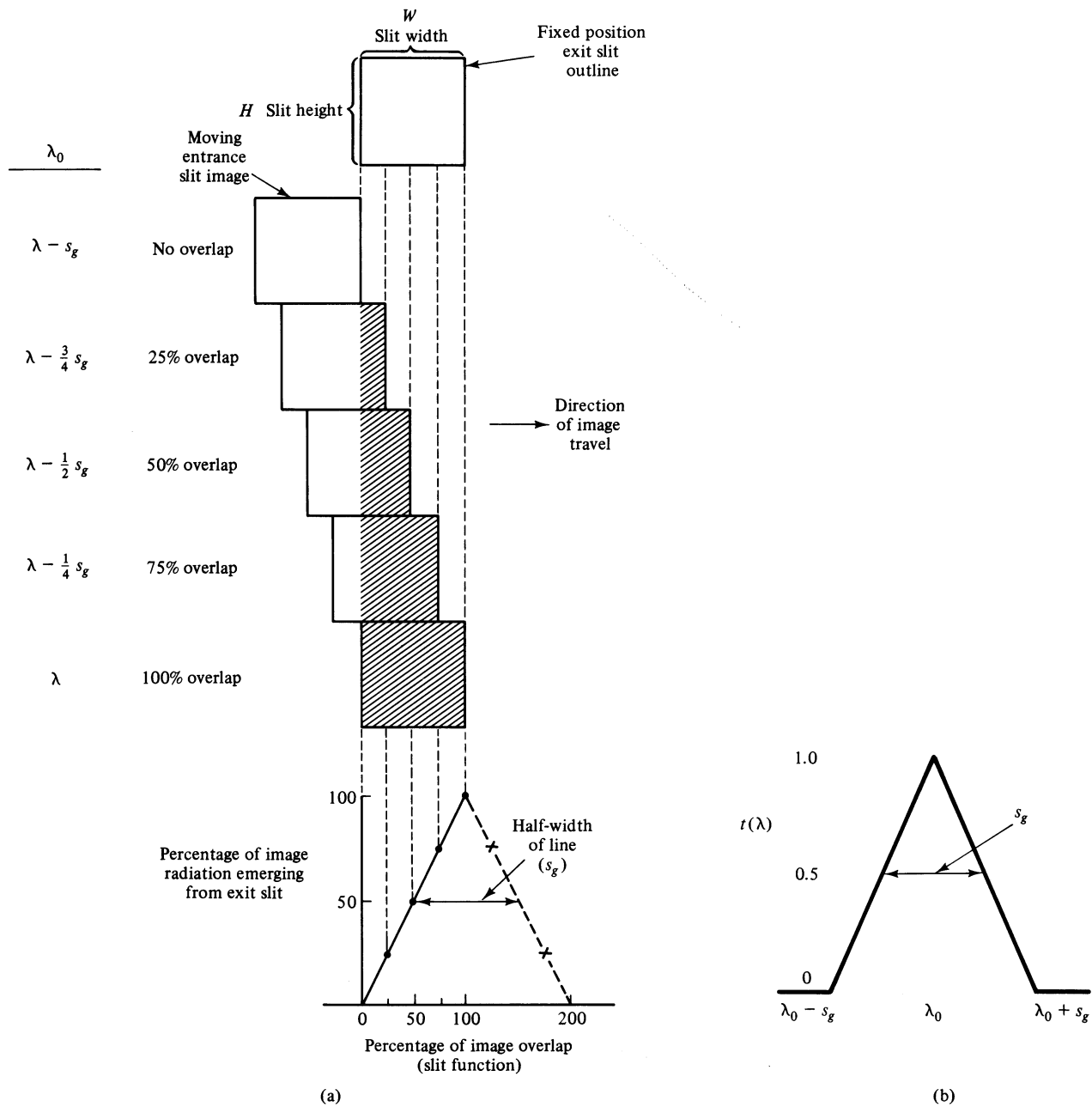
where

R_d is the inverse linear dispersion

W is slit width

In a monochromator, an image of entrance slit is focused at the exit slit:

- when input is **polychromatic**, a monochromated version of the image appears at the exit slit
- when input is **monochromatic** image, rotating the grating angle will **sweep monochromatic image** across the exit slit



The **total intensity** $t(\lambda)$ measured at the exit slit as image is translated is called the **slit function**

- for equal entrance and exit slits, shape is **triangular**
- for unequal entrance and exit slits, shape is **trapezoidal**

with a base of s and half-width of s_g

Mathematically, the slit function is

$$t(\lambda) = 1 - \frac{|\lambda - \lambda_0|}{s_g} \quad \lambda_0 - s_g \leq \lambda \leq \lambda_0 + s_g$$

$$t(\lambda) = 0 \quad \text{elsewhere}$$

where

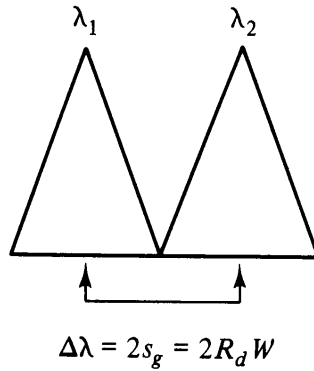
λ_0 is the incident (monochromatic) wavelength at entrance slit

λ_0 is the wavelength setting of the monochromator (the wavelength directed to the center of the exit slit)

Resolution

Resolution quantifies **how well separated** two features are at the exit slit

- closely related to **linear dispersion** (D_l)
(or **angular dispersion** (D_A), and physical dimensions of the monochromator (**through f**))
- slit width W



If the width of a single peak base is s ($= 2s_g$), then two features will just be completely separated when the wavelength difference between them is s

$$s = s = 2s_g = 2R_d W \quad \text{slit - width - limited resolution}$$

Alternatively, we may adjust slit width to obtain resolution of two features separated by s

$$W = \frac{s}{2R_d}$$